

Space Science Activities in China

C. Wang

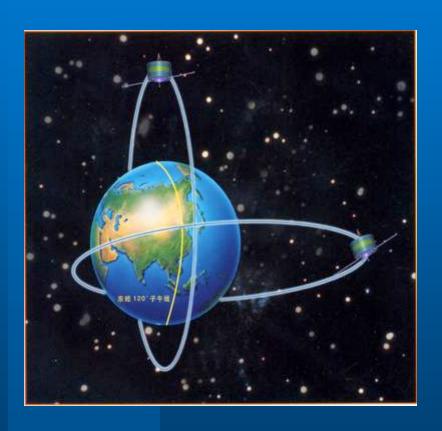
Center for Space Science and Applied Research
Chinese Academy of Sciences

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Outline

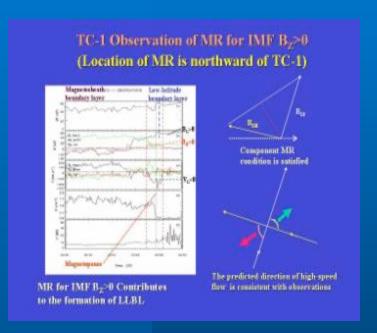
- Recent Advances in Double Star Program
- KuaFu
 - Solar Storm, Aurora and Space Weather Exploration
- SMESE
 - Small Mission on Exploration of Solar Eruption
- Space Physics Exploration Roadmap
- Meridian Project
 - Meridian Space Weather Monitoring Project

Status of DSP



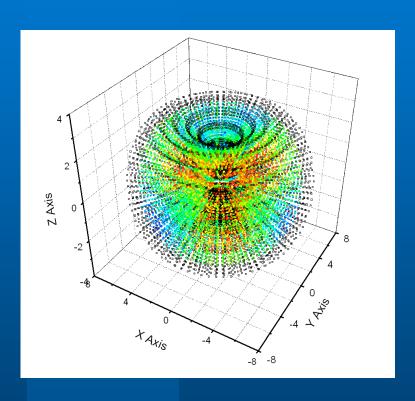
- The DSP mission has been officially extended to the end of 2006.
- By May 2006, TC-1 has obtained a total amount of 175G data, while TC-2 145G.

Double Star Project



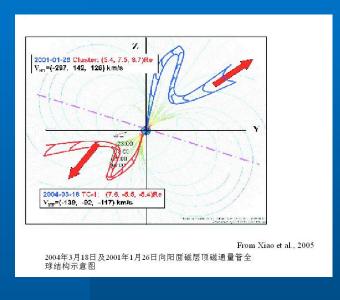
- Data analysis and theoretical study (modeling) related to DSP are very active in China and worldwide.
- First results (24 papers) have been published in *Annales* Geophysice.

Example 1:



 By using of the Image Data from the NUADU/TC2, the distribution of the neutral atom origins has been deduced.

Example 2:



 The occurrence of the magnetic reconnection at the dayside magnetopause under the northward IMF has been observed



KuaFu Solar Storm° Aurora and Space Weather Exploration

L1£polar trip-star project

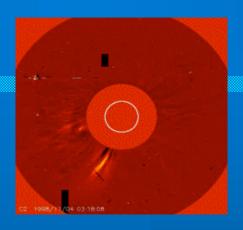
Kuafu-A at L1

Kuafu-B2

Project Leader: Prof. C.-Y. Tu

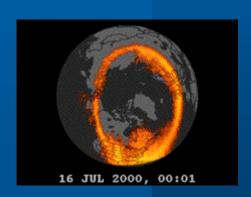
Expected Launch Date: 2012

KuaFu-A at L1



Solar EUV emission
White light CME
Radio wave measurement
local plasma and magnetic field
High energy particles

KuaFu-B1+B2 polar orbit



24 hours Aurora Image, Magnetic field High energy particles

Scientific Objectives

To observe the complete chain of disturbance from the solar atmosphere to the geo-space:

- Solar flares, CMEs,
- Interplanetary clouds, shock waves,
- Their geo-effectiveness, such as sub-storms and magnetic storms, aurora activities

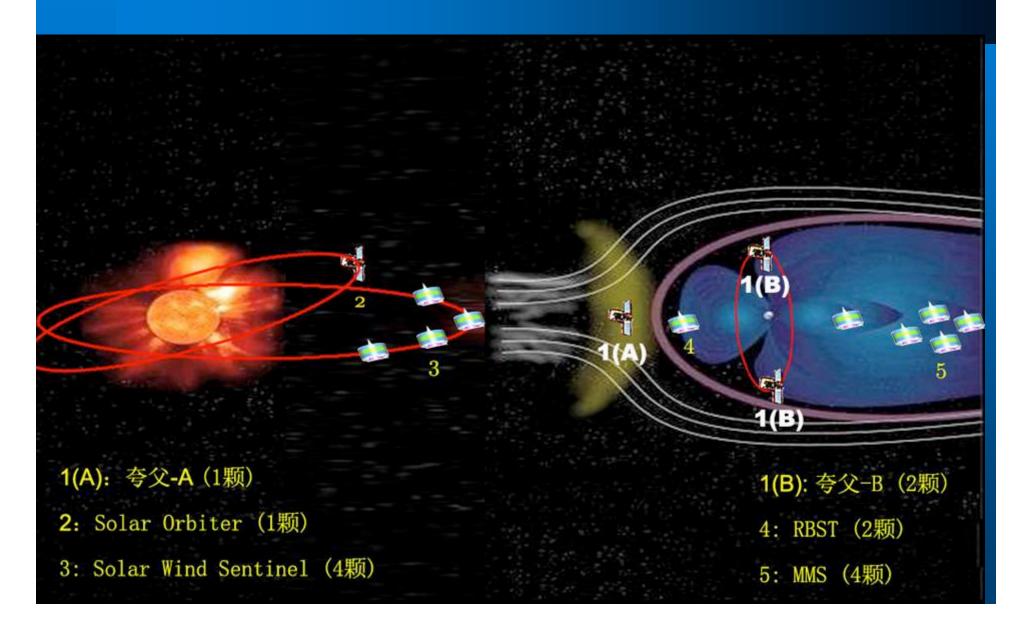
KuaFu-A Payload

Instrument	Mass (kg)	Power (w)	Data Rate (bps)
EUV/FUV Disk Imager	25	20	100k
Coronal Dynamics Imager	35	35	50k
Radio Burst Instrument	12	10	4k
Solar Wind Instrument Package	6	6	4k
Solar Energetic Particle Sensor	8	10	1k
X-ray/ray Detector	3	5	1k
Total	89	86	160

KuaFu-B Payload

Instrument	Mass (kg)	Power (w)	Data Rate (bps)
EUV Aurora Monitoring Camera	21	11	<0.5 M
FUV Aurora Spectrograph	20	4.4	<0.5 M
Wide-Field Aurora Imager	1.5	10	ТВС
Fluxgate Magnetometer	3	4.2	3k
High Energy Particle Experiment	3	4	1k
Neutral Atom Imager	4	4	1k
Total	~60	~50	~0.5 M

Conjunct Observation

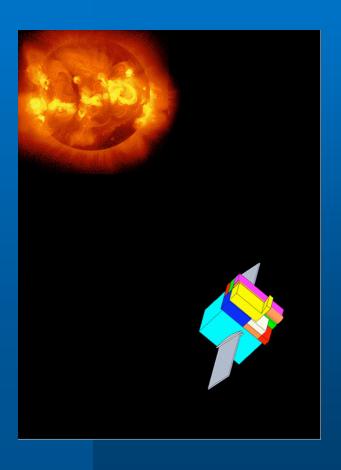


Status

- The Comprehensive Review is under way.
- The KuaFu Promotion Committee have recently been established.

An Brief Introduction to SMESE Small Mission on Exploration of Solar Eruption

France-China Joint Mission

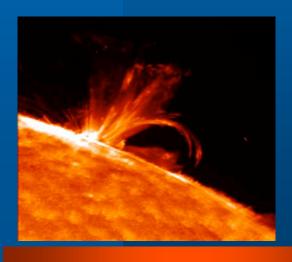


- Purple Mountain Observatory, CAS, China
- Nanjing University, China
- Center for Space Science and Applied Research, CAS, China
- National Astronomical Observatory, CAS, China
- Institut d'Astrophysique Spatiale, CNRS, France
- Observatoire de Paris, France
- Max-Planck-Institute For solar System Research, Germany

Background



The solar corona is the seat of variability which involves basic processes of plasma-magnetic field interaction in the universe.



Explosive energy conversion in the corona has two major types of manifestation:

- coronal mass ejections (CMEs)
- solar flares

Both are basically driven by instabilities of the magnetic field in the corona. 15

SMESE Scientific Objectives

(or the questions to be answered)

- What triggers coronal mass ejections and how are they accelerated in the low corona and out of the corona into the Heliosphere?
- How are CMEs and flares related? More specifically: what is the partition energy between large-scale mass motion, small-scale heating, and particle acceleration?
- How does the Sun accelerate particles to relativistic energies? Low and high energy particles: accelerated by the same mechanism? What are the highest energies to which charged particles can be accelerated in the solar corona?
- How do radiation, conduction and energetic particles participate in the transport of energy in the solar atmosphere?

Brief Description

SMESE:

to observe the solar flares and CMEs for the next Solar Maximum

Main Goals:

- To establish the interconnections between flares and CMEs;
- To follow the disc source region of CME;
- To diagnose the high energy particles accelerated by flares and CMEs;
- To study the energy transportation mechanisms.

all of these are the key ingredients of space weather!

	Wavelength/ Energy	Field of View	Spatial resolution	Time resolution
LYOT: Ly α Imager	121.6 nm	Full Disk	1 arcsec	30s
LYOT: Ly α coronagraph	121.6 nm	>5 solar radii	2.5 arsec	2 min
LYOT: EUV imager	19.5 nm	Full Disk	1 arcsec	30 s
DESIR: Detection of Solar Infra red Radiation	35 and 150 ¶ Ã n	Full Disk	50 arcsec at 35¶Ãn	100 ms
HEBS: HEBS-X	10-500 keV	Full Disk	Full Disk	1s down to 32 ms
HEBS HEBS-G	100 keV- 10 MeV 10 MeV -600 MeV	Full Disk	Full Disk	1s down to 32 ms

	Mass (kg)	Volume (cm ³)	Power (W)	Telemetry
LYOT:	27	58000	39	28 G/day
DESIR	10	20000	10	2 G/day
HEBS	20	23000	18	1 G/day
ECU (computer +mass memory)	4	4000	4	
TM/TC package	2	2000	2	
Structure	12			
Total	75	107000	73	31G/day

SMESE is a small satellite!

Historic role

Missions for solar peak years:

- Cycle 21st: SMM
- Cycle 22nd: YOHKOH
- Cycle 23rd: RHESSI

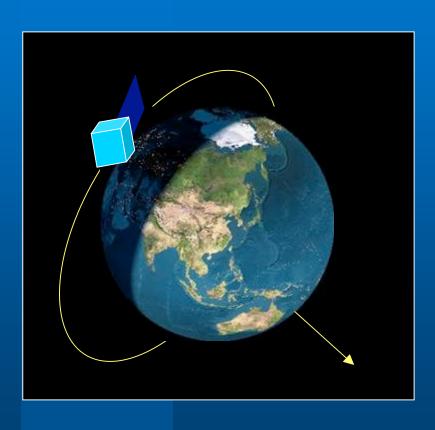


Cycle 24th: SMESE



It may be the unique mission specially for the 24th solar peak years!

Mission Profile



- Mass of payloads: 75 kg
- Power available for the payload: 90- 120W depending on the season
- Size of payloads: 600°600°600mm
- Attitude: Solar-pointed, three-axis stability
- Stability: 0.5 arcsec over 10s
- Telemetry: 31 Gbits/day
- Data recorder: 31 Gbits (solid state)
- Orbit: Polar, Sun Synchronous, 650-750 km
- Launch: 2010
- Lifetime: 2 years minimum
- Communication: X-band

Science management

Science Definition Team (SSDT) has been set up to validate the work of the SMESE consortium during the definition phase. All through the development stages, it ensures that the initial scientific goals are satisfied.

Liaison with ground or space projects.

SSDT membership:

China: Cheng Fang (vice chairman, Nanjing University), Weiqun Gan (Purple

Mountain Observatory),

Chi Wang (CSSAR), Yihua Yan (National Astronomical Observatory)

France: Claude Barat (CESR), Jean Heyvaerts (Strasbourg Observatory), Gérard

Trottet (LESIA),

Jean-Claude Vial (chairman, IAS)

Madhulika Guhathakurta (NASA)

Richard G. Marsden (ESA/ESTEC)

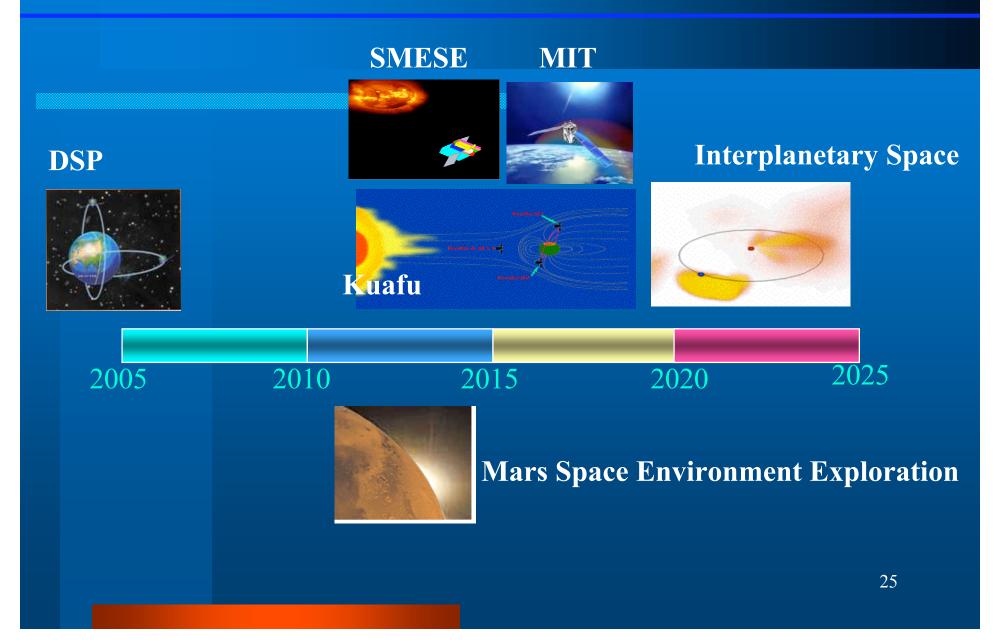
Status

- The phase 0/A Kick Off Meeting has been held in Paris on March 22-24, 2006;
- The Mission Design Review has been held on June 26, 2006. The final report of the review group will be released by the end of August;
- A preliminary requirement review is planned for April 2007;

Status

- On the China side, the preliminary designs of the HEBS instrument and IDHP (Instrument Data Handling Processor) have been funded by CNSA and the support to carry out phase B studies is expected by mid 2007.
- The CNES Scientific Program Committee is expected to give its recommendations regarding the scientific priorities to CNES executive between Oct. 2006 and Feb. 2007.

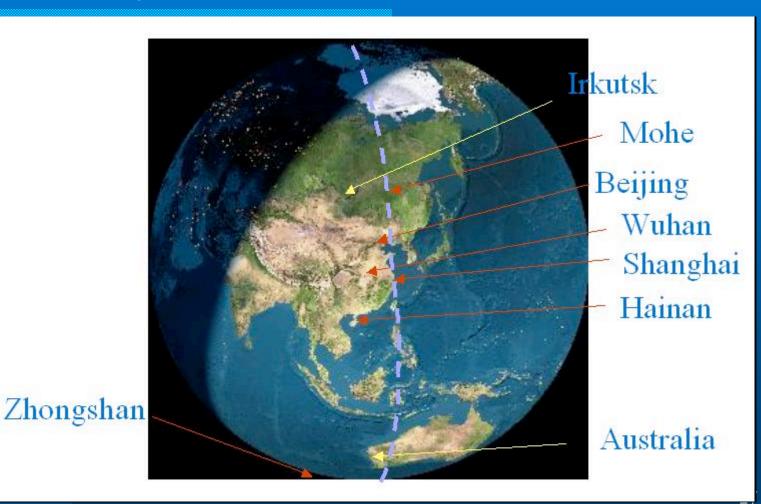
Roadmap for Space Physics Exploration





Meridian Space Weather Monitoring Project (Meridian Project)

It is a Chinese multi-station chain along 120°E to monitor space environment, starting from Mohe, the most northern station in China, through Beijing° Wuhan° Guangzhou and extended to Chinese Zhongshan station in the Antarctic.



Radio Geo-Detection **Magnetic** Subsystem **Subsystem Space Environment Monitoring System Domestic Data Exchange**

Rocket **Optical** Soundin Subsyste Subsyste

Data and Communication System

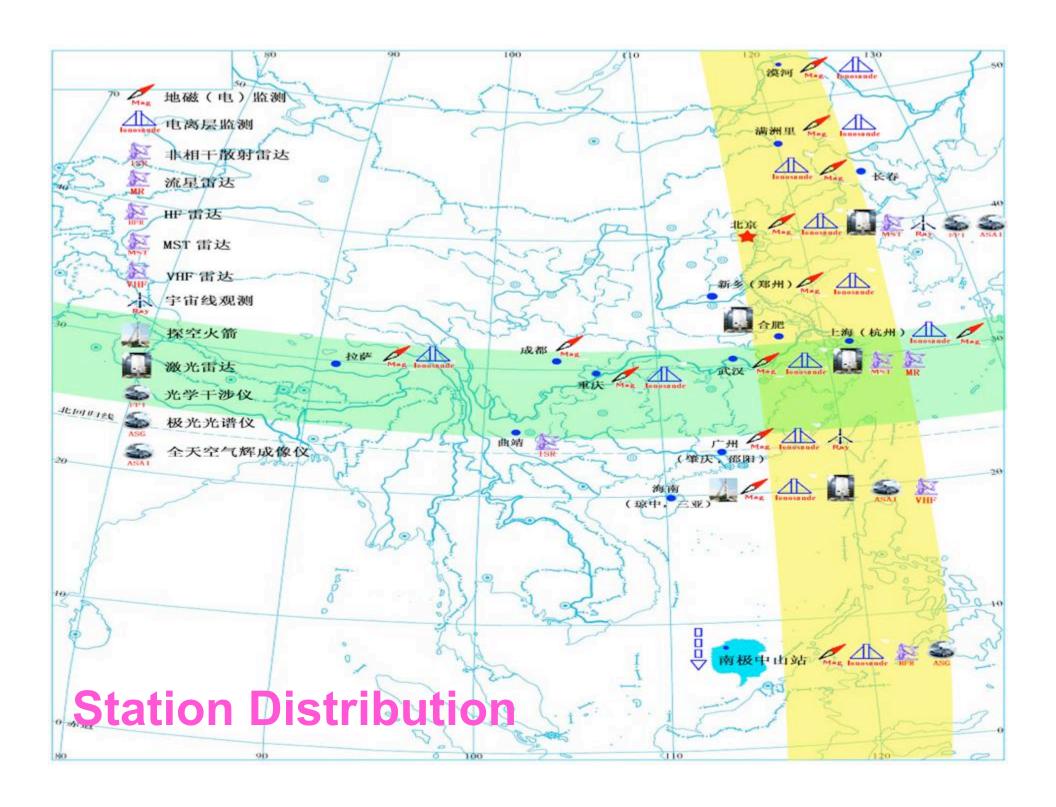
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International **Data Exchange**

Research and Forecast System

Public User

Professional User



No	Station	Lat.	Lon.	Types of Observations
01	Mohe	53.5N	122.4E	Geomagnetic
02	Manzhouli	49.6N	117.4E	Geomagnetic, Ionospheric
03	Harbin	45.5N	126.5E	, lonospheric
04	Changchun	44.0N	125.2E	Geomagnetic
05	Beijing	40.3N	116.2E	Geomagnetic, Ionospheric, Litter, MST radar, Optical atmospheric, IPS, Cosmic rays
06	Xinxiang	34.6N	113.6E	Geomagnetic, Ionospheric
07	Wuhan	30.5N	114.6E	Geomagnetic, Ionospheric, Lidar, MST radar
08	Guangzhou	23.1N	113.3E	Geomagnetic, Ionospheric, Cosmic rays
09	Hainan	19.0N	109.8E	Geomagnetic, Ionospheric, Lidar
10	Zhongshan	69.45	76.4E	Geomagnetic, lonospheric
11	Shanghai	31.1N	121.2E	Geomagnetic
12	Chongqing	29.5N	106.5E	Geomagnetic, Ionospheric
13	Chengdu	31.0N	103.7E	Geomagnetic, Ionospheric
14	Qujing	25.6N	103.8E	Incoherent Scattering Radar (ISR)
15	Lhasa	29.6N	91.0E	Geomagnetic, lonospheric 30

International Collaboration



The International Space Weather Meridian Circle Program (ISWMCP), proposal to connect 120°E and 60°W meridian chains of ground based monitors and enhance the ability of monitoring space environment worldwide.

Status

- The proposal of the Meridian Project has been approved by the China National Development and Reform Commission as a key scientific project in Aug., 2005;
- The feasibility study of the Meridian Project has been evaluated and approved in June, 2006. The total budget is about 22 million US\$;
- The Meridian Project is expected to be finished by the end of 2009.

Summary

- DSP is the first Chinese scientific exploration mission, and has been very successful.
- The comprehensive review of KuaFu project and the phase A study of SMESE are under way.
- Meridian Project has officially approved by Chinese government, and will be finished by 2009.
- International cooperation are encouraged by Chinese government to make joint effort to promote space science missions.

